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## References

- ABRAHAMS, S. C. (1956). Quart. Rev. Chem. Soc., Lond. 10, 407.
- AHMED, F. R. & CRUICKSHANK, D. W. J. (1953). Acta Cryst. 6, 765.
- ALDERMAN, P. R. H. & Owston, P. G. (1956). Nature, 178, 1071.
- BERGHUIS, J., HAANAPPEL, IJ. M., POTTERS, M., LOOP-STRA, B. O., MACGILLAVRY, C. H. & VEENENDAAL, A. L. (1955). Acta Cryst. 8, 478.
- BOOTH, A. D. (1946). Proc. Roy. Soc. A, 188, 77.
- Воотн, А. D. (1947). Proc. Roy. Soc. A, 190, 482.
- BUERGER, M. J. (1937). Z. Kristallogr. 97, 504.

- CAMBI, L. & SZEGÖ, L. (1931). Atti R. Accad. Lincei, 13, 168.
- CRUICKSHANK, D. W. J. (1949). Acta Cryst. 2, 65.
- EWENS, R. G. V. (1948). Nature, 161, 530.
- HOFMANN, K. A. & WIEDE, O. F. (1895). Z. anorg. Chem. 9, 295.
- JAMES, R. W. (1954). The Optical Principles of X-ray Diffraction. Appendix III. London: Bell.
- JENNINGS, J. S. (1939). Ph.D. Thesis, Birmingham.
- JOHANSSON, G. & LIPSCOMB, W. N. (1958). Acta Cryst. 11, 594.
- LEE, D. (1952). PhD. Thesis, Leeds.
- PARKER, H. M. & WHITEHOUSE, W. J. (1932). Phil. Mag. 14, 939.
- PAWEL, O. (1882). Ber. dtsch. chem. Ges. 15, 2600.
- Powel, H. M. & Ewens, R. G. V. (1939). J. Chem. Soc. p. 286.
- REIHLEN, H. & FRIEDOLSHEIM, A. VON (1927). Annalen, 457, 71.
- ROUSSIN, J. (1858). Ann. Chim. Phys. [3], 52, 285.
- SEEL, F. (1942). Z. anorg. Chem. 249, 308.
- VIERVOLL, H. & ÖGRIM, O. (1949). Acta Cryst. 2, 277.
- WILSON, A. J. C. (1949). Acta Cryst. 2, 318.

Acta Cryst. (1958). 11, 604

## Beta-Wolfram Structure of Compounds Between Transition Elements and Aluminum, Gallium and Antimony

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The compounds Nb<sub>3</sub>Al, Nb<sub>3</sub>Ga and Cr<sub>3</sub>Ga have the  $\beta$ -wolfram structure (A15) as determined by the powder method. The space group is  $O_{\hbar}^{3}$ -Pm3n with 2 formula weights in the unit cell. For Nb<sub>3</sub>Al,  $a_{0} = 5 \cdot 187$  Å; for Nb<sub>3</sub>Ga,  $a_{0} = 5 \cdot 171$  Å; for Cr<sub>3</sub>Ga,  $a_{0} = 4 \cdot 645$  Å.

More complete data are given for Mo<sub>3</sub>Ga, V<sub>3</sub>Ga, V<sub>3</sub>Sb and Nb<sub>3</sub>Sb whose preliminary lattice constants were reported previously. For Mo<sub>3</sub>Ga,  $a_0 = 4.943$  Å; for V<sub>3</sub>Ga,  $a_0 = 4.816$  Å; for V<sub>3</sub>Sb,  $a_0 = 4.932$  Å; for Nb<sub>3</sub>Sb,  $a_0 = 5.262$  Å.

The superconducting transition temperature for Nb<sub>3</sub>Al is 17.5 °K.; for Nb<sub>3</sub>Ga, 14.5 °K. and for V<sub>3</sub>Ga, 16.5 °K. The remaining compounds are not superconducting above 1.02 °K.

Three new beta-wolfram compounds have been made in the course of a continuing search for superconductivity in substances with the beta-wolfram structure (Matthias, Geballe, Geller & Corenzwit (1954); Geller, Matthias & Goldstein (1955); Wood & Matthias (1956); Matthias, Wood, Corenzwit & Bala (1956)). These are Nb<sub>3</sub>Al, Nb<sub>3</sub>Ga and Cr<sub>3</sub>Ga. Full data are also given for Mo<sub>3</sub>Ga, V<sub>3</sub>Ga, V<sub>2</sub>Sb and Nb<sub>3</sub>Sb whose preliminary lattice constants were reported by Matthias, Wood, Corenzwit & Bala (1956). Our data for Mo<sub>3</sub>Al which was assigned the beta-wolfram structure in a report of the Climax Molybdenum Co. of Michigan (1951) are also given.

The lattice constants, interatomic distances and superconductivity data are listed in Table 1. The observed intensities are listed in Table 2, together with the intensities calculated from the formula

$$I \propto p |F_{hkl}|^2 \{ (1 + \cos^2 2\theta) / \sin^2 \theta \cos \theta \} \times 10^{-5}$$

where p is the multiplicity factor,  $F_{hkl}$  the structure amplitude, and the remaining term is twice the combined Lorentz and polarization factors. The scattering factors for Nb, Mo, Ga, and Sb were obtained from the paper by Thomas & Umeda (1957), for Cr and Al from the paper by Viervoll & Ögrim (1949) and for V from the International Tables (1935). Dispersion corrections from Dauben & Templeton (1955) were applied.

The agreement between observed and calculated intensities in Table 2 is satisfactory with the exception of the case of Nb<sub>3</sub>Sb which is discussed in the last paragraph of the paper. Absorption most probably accounts for the low intensity of the low angle lines.

The  $\beta$ -wolfram structure belongs to space group

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Table 1. Lattice constants, interatomic distances and superconductivity data for  $A_3B$  compounds

Compound	Lattice constant	Inte	ratomic distanc	ees, A	
	$a_{0}$ Å	B–A	$(A-A)_1$	$(A - A)_{2}$	Superconducting transition temperature °K.
Nb <sub>3</sub> Al	$5 \cdot 187 \pm 0 \cdot 002$	2.900	2.594	3.175	17.5°
Nb <sub>3</sub> Ga	$5 \cdot 171 \pm 0 \cdot 002$	2.890	2.586	3.166	14·5°
Mo <sub>3</sub> Al	$\textbf{4.950} \pm \textbf{0.001}$	2.767	2.475	3.030	Not superconducting above 1.02°
Mo <sub>3</sub> Ga	$4 \cdot 943 \pm 0 \cdot 002$	2.763	2.472	3.026	Not superconducting above 1.02°
V <sub>3</sub> Ğa	$4 \cdot 816 \pm 0 \cdot 002$	$2 \cdot 692$	$2 \cdot 408$	2.948	16·5°
Cr <sub>3</sub> Ga	$4.645 \pm 0.004$	2.596	$2 \cdot 322$	2.844	Not superconducting above 1.02°
$V_3Sb$	$4 \cdot 932 \pm 0 \cdot 002$	2.757	2.466	3.019	Not superconducting above 1.02°
$Mb_3Sb$	$5 \cdot 262 \pm 0 \cdot 002$	2.941	2.631	3.221	Not superconducting above 1.02°

Table 2. Comparison of calculated with observed intensities (visually estimated)

hkl	hkl Nb <sub>3</sub> Al		Nb <sub>3</sub> Ga			MogAl		Mo <sub>3</sub> Ga		V3Ga		Cr <sub>3</sub> Ca		V <sub>3</sub> S⊳		Nb3SP	
	Calc.	Obs.	Calc.	Obs.	Calc.	Obs.	<u>Calc.</u>	Obs.	Calc.	Obs.	Calc.	Obs.	Calc.	Obs.	Calc.	Obs.	
110	12.4	s	2.03	Abs.	11.7	s	2.2	w	0.8	no	0.4	vw	6.1	m	1.8		
200	8.5	s	13.8	5 <sup>a</sup>	7.1	m	12.9	5	1.8	s	2.1	5	4.3	m	27.1	m-s	
210	57.1	vvs	56.4	vvs <sup>b</sup>	50.1	vvs	52.1	vvs	2.2	VS	3.8	vs	2.5	m	60.1	8	
211	18.6	VS	30.9	5	15.7	VS	27.7	vs	3.2	vvs	4.3	vs	8.9	8	59.2	5	
220	1.7	VW	0.3	Abs.	1.5	w	0.3	VVW	0.1	w-m	0.05	VVW	0.9	VW	0.3	Abs.	
310	2.3	w-m	0.5	m-s <sup>b</sup>	2.0	m	0.5	w <sup>a</sup>	0.2	m	0.09	vw	1.6	w	0.4	Abs.	
222	8.4	m-s	6.0	w-m <sup>a</sup>	7.0	VS	5.7	m∼s	0.05	vw	0.3	w	0,1	Abs.	3.5	w	
320	12.1	3	11.8	m	9.9	vs	11.0	5	0.7	m-s	1.8	8	0.7	vw	12.6	m-s	
321	9.1	m-s	14.5	m	7.2	vs	13.1	s	3.6	vvs	6.4	vvs	9.2	8	29.6	s	
400	5.8	m	7.0	w-m	4.7	m-s	6.5	m−s	1.7	8	6.0	vvs	2.9	m	10.2	m-s	
411,330	1.2	vw	0.3	Abs.	1.2	W	0.3	vw <sup>a</sup>	-	-	-	-	7.7	m-s	0.3	vw <sup>a</sup>	
420	2.6	w~m	4.1	vw	2.1	m	3.8	m	-	-	-	-	-	-	8.4	m	
421	11.2	S	11.1	m	9.9	vs	10.8	s	-	-	-	-	-	-	11.8	m-s	
332	2.3	w-m	3.6	vw	1.9	m	3.4	m	-	-	-	-	-	-	7.4	m	
422	0.6	VVW	0.1	Abs.	0.6	vvw	0.2	Abs.	-	-	-	-	-	-	0.1	Abs.	
510,431	1.6	w	0.4	Abs.	1.8	m	0.5	w <sup>a</sup>	-	-	-	-	-	-	0.4	vvw <sup>a</sup>	
520,432	13.0	s	12.9	m	13.4	vvs	14.2	VS	-	-	-	-	-	-	13.0	m-s	
521	3.8	w-m	5.9	w	3.6	m-s	6.4	m-s	-	-	-	-	-	-	11.7	m-s	
440	6.1	m	7.4	w-m	6.4	8	9.1	5	-	-	-	-	-	-	10.2	m-8	
530,433	1.1	vw	0.3	Abs.	1.5	m	0.4	vw	-	-	-	-	-	-	0.2	Abs.	
600,442	2.7	w-m	4.2	vw	3.0	<b>m-</b> 5	5.5	m-s	-	-	-	-	-	-	8.1	m	
610	5.2	m	5.2	w	7.0	VS	7.5	8	-	-	-	-	-	-	5.0	w-m	
611,532	7.1	m	11.2	m	9.2	VS.	16.8	VS	-	-	-	-	-	-	21.3	s	
620	0.8	WVW	0.2	Abs.	1.9	m	0.5	₩	-	-	-	-	-	~	0.2	w	
541	1.9	vw	0.6	Abs.	-	-	-	-	-	-	-	-	-	-	0.4	Abs.	
622	23.1	vs	19.8	m-s	-	-	-	-	-	-	-	-	-	-	6.2	m	
630,542	-	-	-	-	-	-	-	-	-	-	-	-	-	-	34.9	75	
631	~	-		-	-	-	-	-	-	-	-	-	-	-	50.7	vvs	

 $a = \beta$  line coincident; b = Nb line coincident.

The intensities of the various substances are not comparable with each other because of differences in exposure (vs = very strong; s = strong; m = medium; w = weak).

 $O_{\hbar}^{3}$ -Pm3n. The formula is  $A_{3}B$  in which the A atoms occupy position  $6(c): \pm (\frac{1}{4}, \frac{1}{2}, 0; \bigcirc)$ ; the B atoms occupy positions  $2(a): (0, 0, 0; \frac{1}{2}, \frac{1}{2}, \frac{1}{2})$ . Each B atom is coordinated to twelve A atoms at the distance  $B-A(\frac{1}{4}a_{0})/5)$ . Each A atom is coordinated to two Aatoms at the distance  $(A-A)_{1}, (\frac{1}{2}a_{0})$ ; to eight A atoms at distance  $(A-A)_{2}, (\frac{1}{4}a_{0})/6)$ ; and to four B atoms at distance B-A. The structure has been described in the above papers and by Hägg & Schönberg (1954).

The composition of alloys prepared in the arc furnace is not precisely known since losses cannot readily be avoided. Chemical analyses of the samples were not made because more than one phase was present in all cases. In some cases the lattice constant was invariant throughout a series of specimens with different initial proportions of components, indicating constant composition of the phase in question. Where a range of lattice constant was observed, this is reported. When the reactants were sealed in silica tubes, no attempt was made to evacuate the tubes because of the large volume of sample relative to the amount of air in the tube. A rough calculation shows that the amount of oxide or nitride formed would be negligibly small. No nitride or oxide diffraction lines are observed in the films.

X-ray diffraction powder photographs were taken of the resulting product, using a Norelco camera of 114.6 mm. diameter with Cu K radiation for the niobium and molybdenum compounds and Cr K radiation for the vanadium and chromium compounds. Nb<sub>3</sub>Al (Corenzwit, 1958) has a transition temperature to the superconducting state of 17.5 °K. This is the highest transition temperature of any substance known to date except that of Nb<sub>3</sub>Sn (18.05 °K.). It was prepared by heating an aluminum-niobium pressed pellet in a helium arc furnace. The lattice constant ranged from 5.185 Å to 5.194 Å. This niobium-aluminum compound was accompanied by a very small amount of impurity which was probably another niobium-aluminum phase.

Nb<sub>3</sub>Ga was prepared by melting gallium with niobium powder in a sealed fused silica tube at 1200 °C. and subsequently fusing the product in the arc furnace in a helium atmosphere. Without the arc furnace fusion, the  $\beta$ -W structure was not obtained. The specimen with the constants reported in Table 1 was made by the reaction of niobium and gallium in atomic proportions of 2:1, but preferential loss of gallium certainly took place during the arc furnace fusion. The powder pattern showed a small amount of niobium and a very small amount of gallium present. No variation of lattice constant was observed although initial proportions of components were varied from 4:1 to 2:1.

Mo<sub>3</sub>Al was formed by the arc furnace reaction of the stoichiometric amounts of the constituents in a helium atmosphere. Molybdenum was present as an impurity, indicating some aluminum loss in the furnace. The lattice constant was unchanged in a melt containing 58 atomic % Al. Here in addition to the  $\beta$ -W phase and Mo, a third phase was present which was not identified. Mo<sub>3</sub>Al was previously reported by the Climax Molybdenum Co. of Michigan (1951), as having the  $\beta$ -W structure with a lattice constant of 4.951 Å which agrees with the value reported in Table 1.

 $Mo_3Ga$ , unlike Nb<sub>3</sub>Ga, can be formed by reaction in a sealed fused silica tube at 1200 °C. A small amount of free molybdenum was present, indicating some loss of gallium to the walls of the tube. No variation of lattice constant was observed although initial proportions of components were varied from 4:1 to 1:1.

 $V_3Ga$ , was formed by causing the elements to react in stoichiometric proportions in the helium arc furnace. A small amount of a second phase, not yet identified, was present in all cases. No variation in lattice constant was observed when the initial proportions of reactants were varied from 3:1 to 5:1. In one sample made from initial proportions 2V:1 Ga, a third phase was present. The unidentified phases are not VN,  $V_2O_3$ ,  $V_2O_5$ , GaN, nor any of the gallium oxides. The  $V_xN_y$  structures described by Hahn (1949) are not present.

Cr<sub>3</sub>Ga forms in the  $\beta$ -W phase when the components are heated to 1200 °C. in a sealed fused silica tube. When more than the stoichiometric amount of chromium is present, the product is inhomogeneous and may contain crystals with the  $\beta$ -W structure ranging in lattice constant from 4.632-4.649 Å. When the

proportions of chromium and gallium are 2:3 a very interesting structure results. The strongest lines in the powder diagram can be indexed on a primitive cubic cell with  $a_0 = 2.82$  Å. However, weaker lines require a larger unit cell and the true unit cell has not yet been determined. An attempt is now being made to obtain a satisfactory single crystal for further work on this compound.

The two antimony compounds presented the greatest preparative and analytical difficulties of any of the compounds herein reported. They were prepared in sealed fused silica tubes and could not be obtained without the presence of other phases. There appears to be only one other phase accompanying the  $\beta$ -W V<sub>3</sub>Sb phase. Its pattern is closely similar, though not identical, with that of the unidentified phase accompanying the V<sub>3</sub>Ga. Since only one sample of V<sub>3</sub>Sb was available, no information concerning the constancy of the lattice constant could be obtained.

In all of the foregoing cases, the film used for the data in Tables 1 and 2 showed very few relatively weak lines in addition to those of the  $\beta$ -W phase. In the case of Nb<sub>3</sub>Sb there are about 32 extra lines, some of which are of weak to medium strength. Nearly all of the stronger lines can be identified as NbO, primitive cubic,  $a_0 = 4.210$  Å (Brauer (1941) and Andersson & Magnéli (1957)). The remaining lines are not those of Nb, NbO, or any of the forms of  $Nb_2O_5$ , Sb, or  $Sb_2O_3$ . Because of the large number of unidentified lines in this pattern, some of which might be superimposed on those of the Nb<sub>3</sub>Sb compound, the identification of this structure as  $\beta$ -W is less wellestablished. The presence of impurities may account for the discrepancy between the calculated and observed intensity for the  $\{620\}$  reflection.

## References

- ANDERSSON, G. & MAGNÉLI, A. (1957). Acta Chem. Scand. 11, 1065.
- BRAUER, G. (1941). Z. anorg. Chemie, 248, 1.
- Climax Molybdenum Co. of Michigan (1951). Second Annual Report Arc-Cast Molybdenum-Base Alloys, NR 034-401.
- CORENZWIT, E. (1958). J. Phys. Chem. Solids. To be published.
- DAUBEN, C. H. & TEMPLETON, D. H. (1955). Acta Cryst. 8, 841.
- GELLER, S., MATTHIAS, B. T. & GOLDSTEIN, R. (1955). J. Amer. Chem. Soc. 77, 1502.
- HÄGG, G. & SCHÖNBERG, N. (1954). Acta Cryst. 7, 351.
- HAHN, H. (1949). Z. anorg. Chemie, 258, 58.
- International Tables for the Determination of Crystal Structures (1935). Berlin: Borntraeger.
- MATTHIAS, B. T., GEBALLE, T., GELLER, S. & CORENZWIT, E. (1954). *Phys. Rev.* 95, 1435.
- MATTHIAS, B. T., WOOD, E. A., CORENZWIT, E. & BALA, V. B. (1956). J. Phys. Chem. Solids, 1, 188.
- THOMAS, L. H. & UMEDA, K. (1957). J. Chem. Phys. 26, 293.
- VIERVOLL, H. & ÖGRIM, O. (1949). Acta Cryst. 2, 277.
- WOOD, E. A. & MATTHIAS, B. T. (1956). Acta Cryst. 9, 534.

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